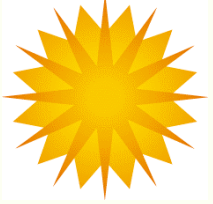


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Energy In Our Daily Lives

Our Energy Sources, Budgets,
Expenditures

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Where Does Energy Come From

- Ultimately, from the Big Bang
 - Energy is, after all, conserved
- In our daily lives: 93% Sun, 7% nuclear
 - Food energy: sunlight, photosynthesis
 - Hydroelectric energy: sunlight-driven water cycle (7%)
 - Fossil Fuels: Stored deposits of plant energy (85%)
 - Wind Energy: solar-driven weather (< 1%)
 - Solar Energy: well...from the sun, of course (< 1%)
 - Our nuclear energy, in essence, derives from products of former stars (supernovae)

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World Energy Budget (annually)

| Source | 10^{18} Joules/yr | Percent of Total |
|-------------------|---------------------|--------------------|
| Petroleum | 158 | 40.0 |
| Coal | 92 | 23.2 |
| Natural Gas | 89 | 22.5 |
| Hydroelectric | 28.7 | 7.2 |
| Nuclear Energy | 26 | 6.6 |
| Biomass (burning) | 1.6 | 0.4 |
| Geothermal | 0.5 | 0.13 |
| Wind | 0.13 | 0.03 |
| Solar Direct | 0.03 | 0.008 |
| Sun Abs. by Earth | 2,000,000 | then radiated away |

circa 2000

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Where does the sun get its energy?

- Thermonuclear fusion reactions in the sun's center
 - Sun is 16 million degrees Celsius in center
 - Enough energy to ram protons together (despite mutual repulsion) and make deuterium, then helium
 - Reaction per mole 20 million times more energetic than chemical reactions, in general

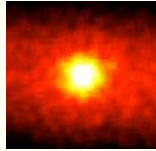
4 protons:
mass = 4.029

^4He nucleus:
mass = 4.0015

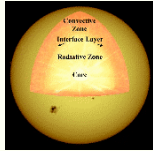
+ 2 neutrinos, photons (light)

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$E = mc^2$ in Sun



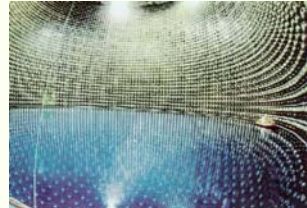
- Helium nucleus is *lighter* than the four protons!
- Mass difference is $4.029 - 4.0015 = 0.0276$ a.m.u.
 - 1 a.m.u. (atomic mass unit) is 1.6605×10^{-27} kg
 - difference of 4.58×10^{-29} kg
 - multiply by c^2 to get 4.12×10^{-12} J
 - 1 mole (6.022×10^{23} particles) of protons $\rightarrow 2.5 \times 10^{12}$ J
 - typical chemical reactions are 100-200 kJ/mole
 - **nuclear fusion is ~20 million times more potent stuff!**

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Solar Energy Output Forms

- 2% in neutrinos: very light, non-interactive
 - more than ten billion per second course through your fingernail
 - fly through earth, as if it weren't even there
 - detected in rare interaction events in huge underground detectors



“Super-K” underground neutrino detector in Japan,
half full of water

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Solar Energy Output Forms, continued

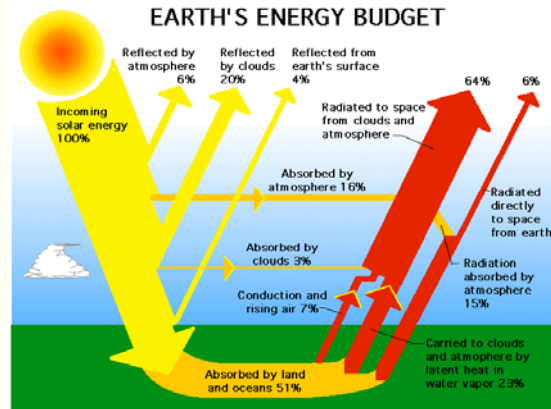
- 98% in light: photons
 - Each photon takes about a million years to clear the annoying electrons in solar plasma
 - 8 minutes once free to reach earth
- 1370 Watts per square meter incident light power
 - Most makes it through atmosphere and reaches us here
 - That which is not reflected is re-radiated back to space
 - after warming us up
 - Hugely abundant: don't have to drill or mine for it

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Where does the sunlight go?

EARTH'S ENERGY BUDGET





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Human Energy Requirements

- 1,500 Calories per day just to be a couch-potato
 - 6,280,000 J
- Average human power consumption is then:
 - 6.28 MJ / 86,400 seconds \approx 75 W
 - We're like light bulbs, constantly putting out heat
- Need more like 2,000 Cal for active lifestyle
 - 100 W of power

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Energy from Food

| Nutrition Facts | |
|---|----------------------|
| Serving Size 2/3 cup (52g) Servings Per Container: 6 | |
| Amount Per Serving | % Daily Value* |
| Calories 180 | Calories from Fat 60 |
| Total Fat 7g | 10% |
| Saturated Fat 1g | 4% |
| Cholesterol 0mg | 0% |
| Sodium 60mg | 3% |
| Potassium 230mg | 7% |
| Total Carbohydrate 28g | 9% |
| Dietary Fiber 7g | 27% |
| Sugars 12g | |
| Protein 10g (6g from soy) | |
| Vitamin A 0% | Vitamin C 2% |
| Calcium 6% | Iron 15% |

- Energy from fat, carbohydrates, protein
 - 9 Calories per gram for fat
 - 4 Calories per gram for carbohydrate
 - Fiber part doesn't count
 - 4 Calories per gram for protein
- Calculate 63 fat, 84 CH, 40 protein Cals
 - total is 187 Calories (180 is in the ballpark)
- 1 Calorie (kilo-calorie) is 4,184 J
 - 180 Cal = 753 kJ
 - set equal to $mgh \rightarrow$ climb 1100 m vertically, assuming perfect efficiency

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Not So Fast...

- Human body isn't 100% efficient: more like 25%
 - To put out 100 J of mechanical work, must eat 400 J
 - 180 Calorie candy bar only gets us 275 m, not 1100 m
- Maximum sustained power output (rowing, cycling) is about 150-200 W (for 70 kg person)
 - Consuming 600-800 W total, mostly as wasted heat
 - For 30 minutes $\rightarrow 800 \text{ J/s} \times 1800 \text{ s} = 1.44 \text{ MJ} = 343 \text{ Cal}$
- Can burst 700 W to 1000 W for < 30 sec
 - put out a full horsepower momentarily!

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Most impressive display of human power

- The Gossamer Albatross crossed the English Channel in 1979, powered by Bryan Allen
 - Flight took 49 minutes, wiped Bryan out!
 - Sustained power out \sim 250 W



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Aside: Human mass balance

- No nuclear power in our stomachs, so mass is conserved
 - mass in = mass out, assuming constant weight
 - burning Calories \neq losing weight, not directly, anyway
- Breathing: an important element in mass balance
 - lose about a pound per day through nose/mouth!
 - breathe in O_2 , breathe out CO_2 : donating carbon to air
 - breathe in dry air, exhale moist air (H_2O loss)
- Trees get their mass through inverse process

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Human Energy Requirements Summarized




- We need chemical energy from food to run
 - Ultimate source is sun, long chain of events to twinkies
 - Constantly burn energy at rate of 75-100W
 - We spend energy at about 25% efficiency
 - Maximum sustained power is 150-200 W
 - actually burn 4 times this due to inefficiencies

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Chemical Energy: Gasoline


- Gasoline and other combustibles are about as energy-rich as the fat we eat: 11 Calories/gram
 - Jet fuel, crude oil, kerosene, you name it



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Fuel Efficiency



- Can calculate miles-per-gallon based on this info:
 - 30 m/s requires 50 kW to fight air drag (Lecture 8)
 - Go one mile in 54 seconds at this speed (67 m.p.h.)
 - $50 \text{ kW} \times 54 \text{ seconds} = 2.68 \text{ MJ} = 640 \text{ Calories}$
 - Assuming 30% engine efficiency (lots of heat), need $640 \times 3.3 = 2100 \text{ Calories}$, or 192 grams of fuel
 - One gallon is 3.5 kg \rightarrow ~20 miles-per-gallon!
- Improvement via aerodynamic drag reduction
 - also helps to go slower (v^2 dependence)

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Energy Expenditure


- Per capita energy production in U.S. at > 10 kW
 - times 86,400 seconds per day is about 1 GJ per day!
 - 1,000,000,000 J per day per person
 - 250,000 Calories
 - Demands 23 kg (6 gallons) of gas per day per person
 - Or equivalently 38 kg (85 lb) of coal (at 6 Cal/gram)
- Most of this expenditure is industrial
 - Production of consumer goods
- Most residential/commercial energy used for heat

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Solar Alternative



- Once fossil fuels are exhausted (coming soon!), need alternative production source
- Straight to solar may be smart
- 1370 W/m^2 incident on earth, 900 W/m^2 typically available to ground panel in full sun
 - take day/night and clouds into consideration: 200 W/m^2 average
 - silicon photovoltaics about 15% efficient $\rightarrow 30 \text{ W/m}^2$
- Each person would need 300 square meters of panels to cover *all* of our nation's energy needs
 - for just our electricity needs, would need square in desert 100 miles on a side \rightarrow not impossible!

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References and Assignments

- References
 - *Energy and the Environment*, Rinstinen & Kraushaar
 - *Energy*, by Gordon Aubrecht, Prentice Hall, 1995
 - *Energy: A Guidebook*, by Janet Ramage (British)
- Course on subject: Physics 12: Energy & Environ.
 - Spring Quarters (I'll teach Spring 2009)
- Midterm Reviews:
 - Wed. 4/30 6:30 PM to 8:20 PM; Pepper Canyon 122 (Tom)
 - Thu. 5/01 8:00 PM to 9:50 PM; Center 212 (Jim)
- Scantron form # 101864-PAR-L & No. 2 pencil
- Assignments:
 - HW for 5/09: Hewitt 7.E.42, 7.P.9, 6.R.16, 6.R.19, 6.R.22, 6.R.23, 6.E.8, 6.E.12, 6.E.43, 6.P.6, 6.P.12, 8.R.29, 8.E.47, 8.P.9

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