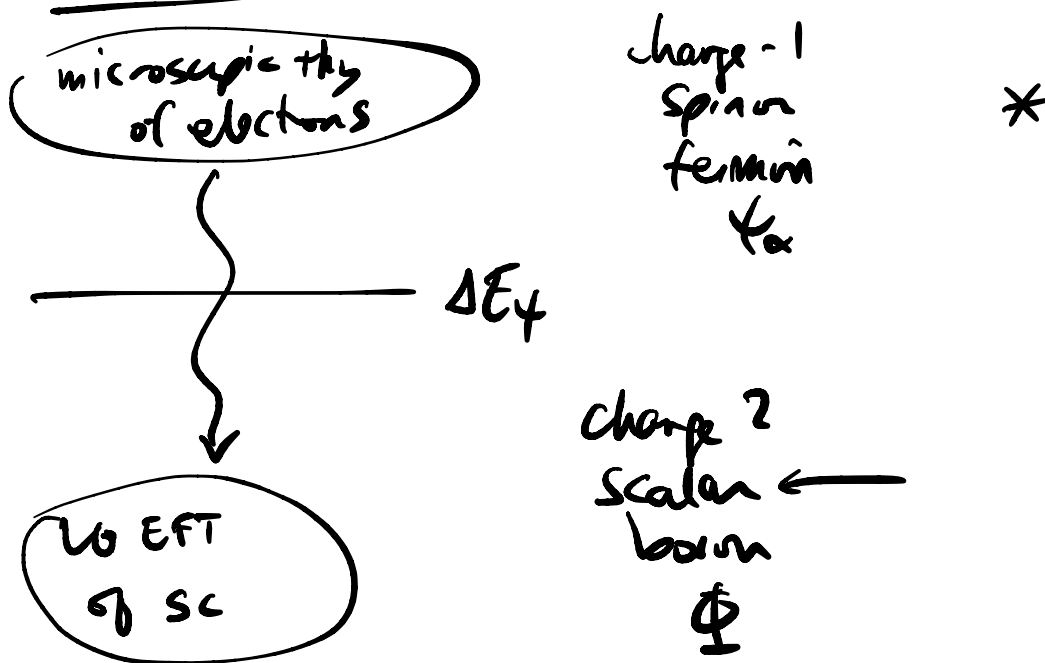


Superconductors & Superfluids, cont'd



Simple case (s-wave): $\Phi \sim \psi_\uparrow \psi_\downarrow \in \text{exp} = \# \psi_\uparrow \psi_\downarrow$ Cooper pair.

eg: $S[\psi] = S_2[\psi] + \int dt d^d x u \psi_{\alpha}^{\dagger} \psi_{\alpha} \psi_{\beta}^{\dagger} \psi_{\beta} + \text{h.c.}$

$u > 0$ repulsive

$\Rightarrow S_2[\psi] = \int dt \int d^d k \psi_k^{\dagger} (i\partial_t - \epsilon(k)) \psi_k$

Mean Field Theory

manover: $S_{\text{MFT}}[\psi] = S_2[\psi] - \int dt d^d x u \langle \psi \psi \rangle \psi^{\dagger} \psi + \text{h.c.}$

$= S_2[\psi] - \int dt d^d x u \bar{\Phi} \psi^{\dagger} \psi + \text{h.c.}$

if $\Phi \neq 0$ ψ gets a mass.

$$\Rightarrow \Delta E_\psi = u \langle \Phi \rangle$$

Why might this work?

Hubbard-Stratonovich transform:

(ep1) in $D=0+0$

$$Z = \int_{-\infty}^{\infty} dx e^{iS(x)}$$

$S \ni x^4$

$$e^{-iux^4} = \frac{1}{\sqrt{\pi u i}} \int_{-\infty}^{\infty} d\sigma e^{-\frac{\sigma^2}{i2} - 2ix^2\sigma}$$

(ep2) in $D=0+1$

$$Z = \int_{\mathcal{C}} d^2x e^{iS(x, \bar{x})}$$

$$e^{-iux^2\bar{x}^2} = \frac{1}{i\pi u} \int_{\mathcal{C}} d^2\sigma e^{-\frac{|\sigma|^2}{i4} - ix^2\sigma + i\bar{x}^2\sigma}$$

$x^2 \rightsquigarrow \psi\psi$

$$Z = \int [\mathcal{D}\psi \mathcal{D}\bar{\psi}] e^{iS[\psi]}$$

$$= \# \int [\mathcal{D}\psi \mathcal{D}\bar{\psi} \mathcal{D}\sigma \mathcal{D}\bar{\sigma}] e^{iS_2[\psi] + i \int d^D x \bar{\sigma} \psi \uparrow \psi \downarrow + h.c.}$$

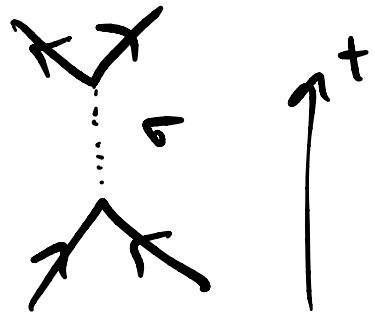
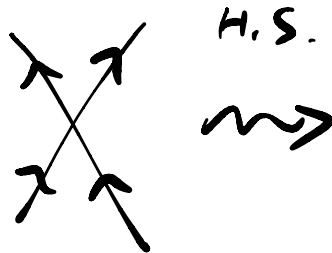
we "decoupled the ψ^4 term". $e^{-\int d^D x \frac{|\sigma|^2}{i\mu}}$

notice: At the saddle of σ integral,

$$\underline{\sigma} = \mu \psi \uparrow \psi \downarrow.$$

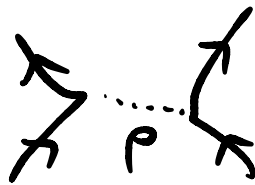
Note that we made a choice to decouple in the "BCS channel"

vs "density channel"



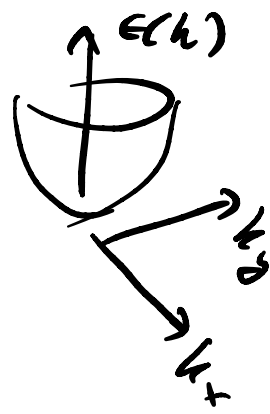
↘ H.S'

$$S(p, \psi) = \int (p \psi^\dagger \psi + \frac{p^2}{2\mu})$$



$\int D\psi$ is now gaussian.

let's choose $\epsilon(k) = \frac{k^2}{2m} - \mu$



$$I_4[\sigma] \equiv \int D\psi D\bar{\psi} e^{i \int dt d^d x L}$$

$$L = \psi^\dagger (i\partial_t - \epsilon(i\nabla)) \psi + \bar{\sigma} \psi_\uparrow \psi_\downarrow + \sigma \psi_\uparrow^\dagger \psi_\downarrow^\dagger$$

$$= (\psi^\dagger \quad \psi) \underbrace{\begin{pmatrix} i\partial_t - \epsilon(i\nabla) & \sigma \\ \bar{\sigma} & -(i\partial_t - \epsilon(-i\nabla)) \end{pmatrix}}_M \begin{pmatrix} \psi \\ \psi^\dagger \end{pmatrix}$$

$$I_4[\sigma] = \det(M \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}) = \det M$$

$$= e^{\text{Tr} \log M}$$

If σ is constant in space:

$$= e^{\int d^3 k \log(\omega^2 - \epsilon_k^2 - |\sigma|^2) + O(\partial\sigma)}$$

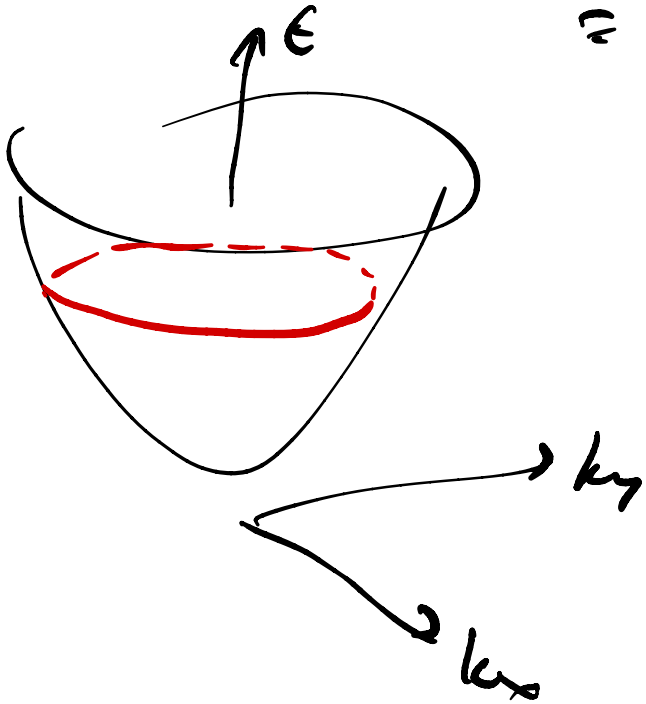
$$Z = \int [D\sigma D\bar{\sigma}] e^{-\int d^D x \frac{|\sigma|^2}{2u} + \int d^D k \log(\omega^2 - \epsilon_k^2 - |\sigma|^2)} + O(\partial\sigma)$$

Sometimes we can do this integral by saddle point

① one example: when the Fermi surface

$$= \{k \mid \epsilon(k) = 0\}$$

$$\epsilon(k) = \frac{k^2}{2m} - \mu$$



② large N # of fermion species

$$\alpha = \uparrow, \downarrow$$

$$\rightsquigarrow \alpha = 1 \dots N$$

$$D = \frac{\delta(\text{exponent})}{\delta \bar{\sigma}} = i \frac{\sigma}{2u} + \int d^D k \frac{2\sigma}{\omega^2 - \epsilon_k^2 - |\sigma|^2 + i\epsilon}$$

do $\int d^D k$ by residues

$$= i \frac{\sigma}{2u} + \int d^D k \frac{2\sigma i}{2\sqrt{\epsilon_k^2 + |\sigma|^2}}$$

→ Gap equation:

$$1 = -2u \int d^d p \frac{1}{\sqrt{\epsilon(p)^2 + |\underline{\sigma}|^2}} \quad (*)$$

value of $\underline{\sigma}$ determines ΔE_F .

Note: • a solution of * requires $u < 0$.

• ^{IF} $\Lambda^d u \rightsquigarrow u(p, p')$

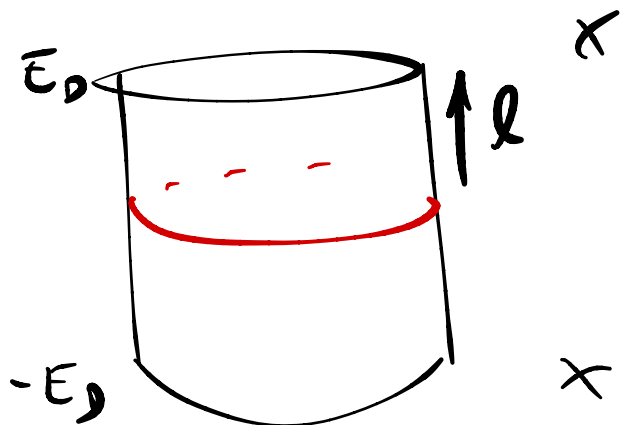
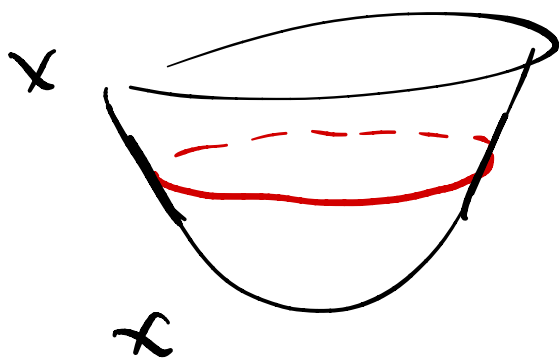
$$\sigma(\vec{p}) = -\frac{1}{2} \int d^d p' \frac{u(p, p') \sigma(p')}{\sqrt{\epsilon(p')^2 + |\sigma(p')|^2}}$$

• To do the integral: expand about the FS

$$\epsilon(p) \approx v_F l \equiv \epsilon$$

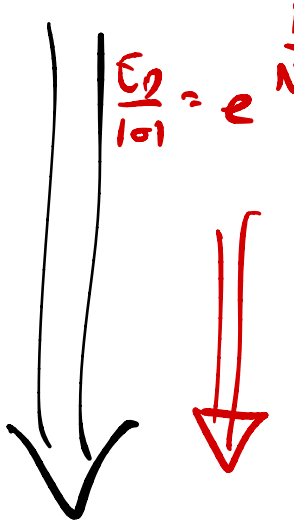
$$v_F = \left. \frac{\partial \epsilon}{\partial p} \right|_{p_F}$$

$$l \equiv p - p_F$$



$$\begin{aligned} \Rightarrow 1 &= -2u \int \frac{d^d p}{\sqrt{\epsilon(p)^2 + |\sigma|^2}} \\ &\approx -2u \int_{FS} \frac{d^{d-1} k}{v_F} \int_{-\epsilon_0}^{\epsilon_0} \frac{d\epsilon}{\sqrt{\epsilon^2 + |\sigma|^2}} \\ &= N|u| \log \left(\frac{\epsilon_0 + \sqrt{\epsilon_0^2 + |\sigma|^2}}{|\sigma|} \right) \approx N|u| \log \frac{\epsilon_0}{|\sigma|} \end{aligned}$$

$$\begin{aligned} N &\equiv \int_{FS} \frac{d^{d-1} k}{v_F} \quad \epsilon_0 \gg |\sigma| \\ &= \text{d.o.s at the FS} \end{aligned}$$

$\frac{\epsilon_0}{|\sigma|} = e^{\frac{1}{N|u|}}$


$$|\sigma| = \frac{2\epsilon_0}{e^{\frac{1}{2N|u|}} - 1} \quad \text{IF } Nu \ll 1 \approx 2\epsilon_0 e^{-\frac{1}{N|u|}}$$

- non-perturbative in u !!
- generates a hierarchy of scales.

Non-relativistic scalar fields:

$$S_{LG}[\Phi] = \int d^d x dt \left(\underbrace{\Phi^* i \partial_t \Phi}_{\text{cancel}} - \frac{\vec{\nabla} \Phi \cdot \vec{\nabla} \Phi}{2m} - V(|\Phi|) \right)$$

~~$+ a \Phi^* \Phi$~~
 is an irrelevant perturbation.

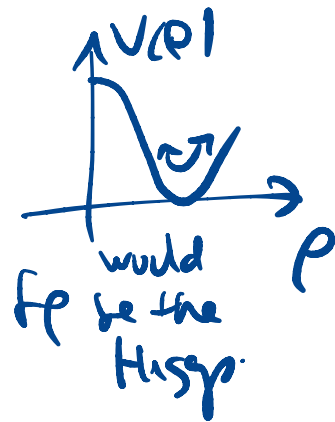
Number & Phase angle: $\frac{\partial L}{\partial \Phi} = \Phi^* (i \partial_t - \dots)$

In terms of $\Phi = \sqrt{\rho} e^{i\psi}$

$$L = \underbrace{\frac{i}{2} \partial_t \rho}_{\text{total deriv.}} - \rho \partial_t \psi - \frac{1}{2m} (\rho (\nabla \psi)^2) + \frac{1}{4\rho} (\nabla \rho)^2 - V(\rho)$$

$$[\rho(x,t), \psi(y,t)] = i \int d^d (\vec{x} - \vec{y})$$

\uparrow number
 $\int d^d x \rho = \# \text{ of particles.}$
 \uparrow phase of condensate.



⇒ No Higgs mode in vanilla
NR SF $\omega < \text{SC}$.

Superfluids: choose $V(\rho) = g^2 (\rho - \bar{\rho})^2$

expand $\sqrt{\rho} = \sqrt{\bar{\rho}} + h$. $h \ll \sqrt{\bar{\rho}}$

$$\Rightarrow L = -2\sqrt{\bar{\rho}} h \partial_t \psi - \frac{\bar{\rho}}{2m} (\nabla \psi)^2 - \frac{1}{2m} (\vec{\nabla} h)^2 - 4g^2 \bar{\rho} h^2 + \mathcal{O}(h^3)$$

Do $\int \rho h$

$$\rightsquigarrow \bar{\rho} \partial_t \psi - \frac{1}{4g^2 \bar{\rho} - \frac{\bar{\rho}^2}{2m}} \partial_t \psi - \frac{\bar{\rho}}{2m} (\vec{\nabla} \psi)^2$$

$$= \frac{1}{4g^2} (\partial_t \psi)^2 - \frac{\bar{\rho}}{2m} (\nabla \psi)^2 + \dots \quad k \ll \sqrt{8g^2 \bar{\rho} m}$$

TRUE
LOW-E
EFT of SF

$$= \frac{\bar{\rho}}{2m} \partial_\mu \psi \partial_\mu \psi + \dots$$

X γ model.

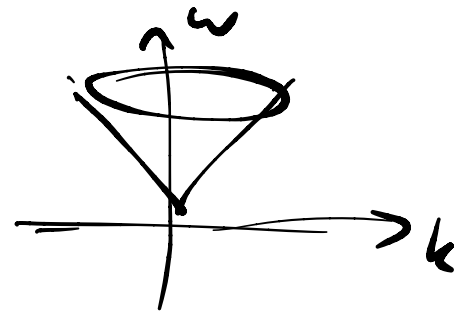
in the
right units
for t, x

$$\Rightarrow \omega^2 = \frac{2g^2 \bar{\rho}}{m} k^2.$$

[Bogoliubov]

relativistic dispersion

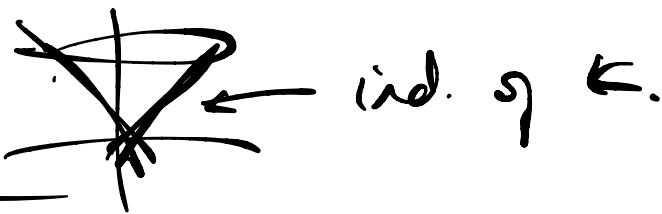
$$v_s = g \sqrt{\frac{2\bar{\rho}}{m}}$$



v_s : Relativistic theory

$$V(\Phi) = g|\Phi|^4 + \dots$$

$$L = \partial_\mu \Phi^* \partial^\mu \Phi - \kappa (\Phi^* \Phi - v^2)^2$$



Superflow: If you start a current, it keeps going.

Whence: $\vec{j} = \frac{\rho}{m} \vec{\nabla} \varphi$ and $\varphi \cong \varphi + 2\pi$.

Consider $x \cong x + L$. set up a current J_x .

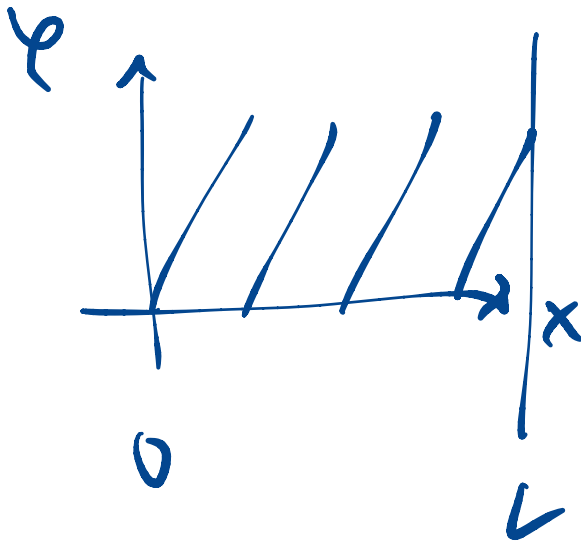
Requires $\varphi(x) = mvx \Rightarrow J_x = \rho v$

$$\left. \begin{array}{l} \psi \approx \psi + 2\pi \\ \text{and} \\ x \approx x + L \end{array} \right\} \Rightarrow$$

$m v L \in 2\pi \hbar$
is quantized

$$\hbar \Rightarrow h$$

$\frac{m v L}{2\pi} \equiv$ vorticity
of Φ

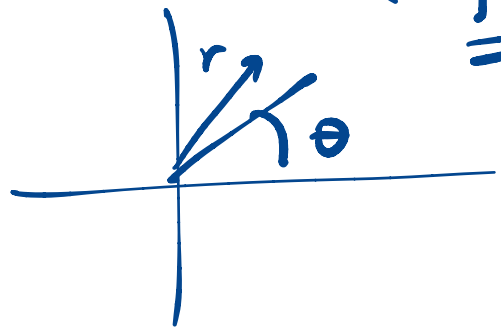


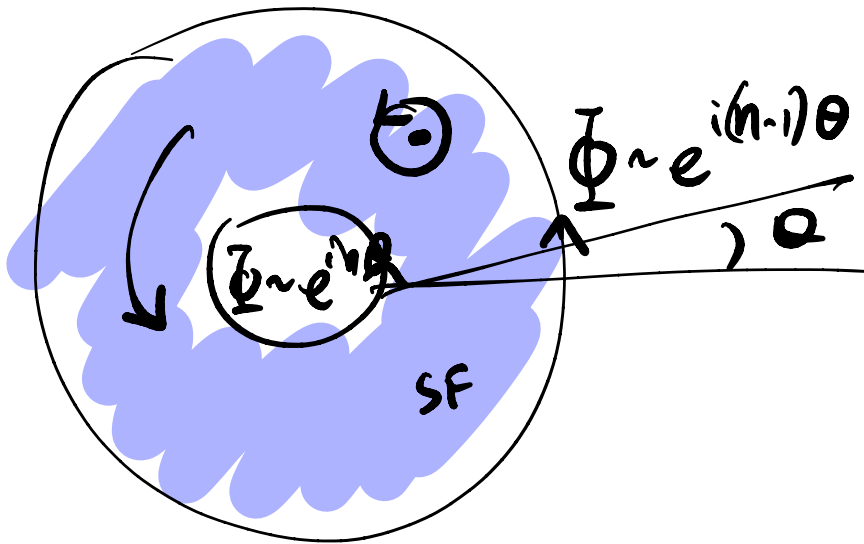
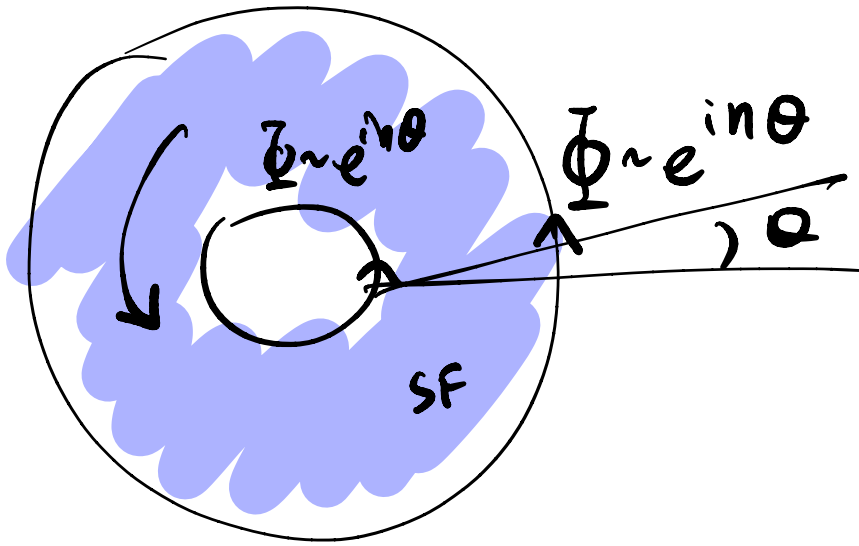
how can h
change?

$$\Phi = \underbrace{|\Phi_0|}_{\sim} e^{i\psi}$$

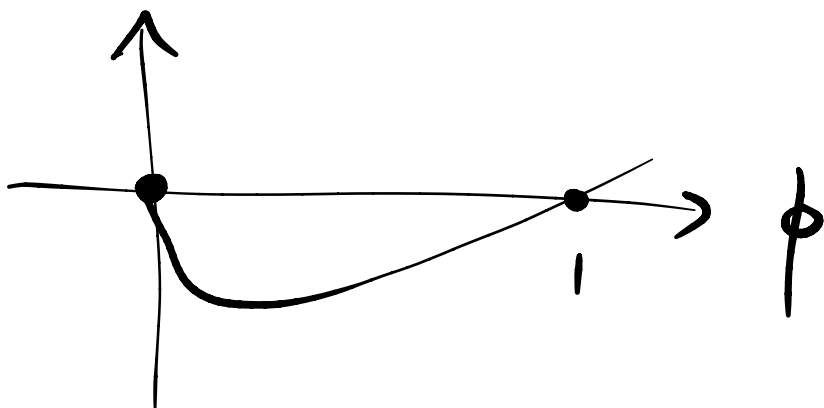
Can only change by vortices

= locus where $\Phi(r, \theta) \sim r e^{i\theta}$
ie $\psi \sim \theta$





$$E(\phi^2) \stackrel{\phi^2 \ll \dots}{\sim} \phi^2 \ln \phi$$



Spectrum of \mathcal{Q} 's

27

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-27

$\mathcal{Q} + 1$

$\mathcal{Q} + 1$

\mathcal{Q}

\mathcal{Q}

\mathcal{C}

