


UCSD Physics 12



Forms of Energy II

Wind, Chemical, Food, Mass-energy, Light

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The Physics 12 Formula List

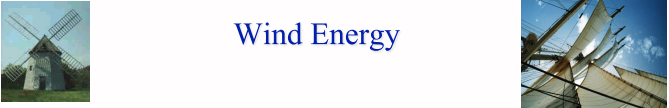
- Lots of forms of energy coming fast and furious, but to put it in perspective, here's a list of formulas that you'll need to use:

Relation Type	Formula
Work as force times distance	$W = F \cdot d$
Kinetic Energy	$K.E. = \frac{1}{2}mv^2$
(Grav.) Potential Energy	$E = mgh$
Heat Content	$\Delta E = c_p m \Delta T$
Power	$P = \Delta E / \Delta t$
Mass-energy	$E = mc^2$
Radiative Flux	$F = \sigma T^4$

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Wind Energy

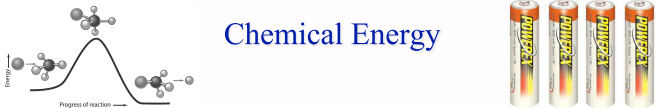


- Wind can be used as a source of energy (windmills, sailing ships, etc.)
- Really just kinetic energy
- Example: wind passing through a square meter at 8 meters per second
 - Each second we have 8 cubic meters
 - Air has density of 1.3 kg/m³, so (8 m³)×(1.3 kg/m³) = 10.4 kg of air each second
 - $\frac{1}{2}mv^2 = \frac{1}{2} \times (10.4 \text{ kg}) \times (8 \text{ m/s})^2 = 333 \text{ J}$
 - 333 J every second → 333 Watts of available power per square meter (but to get *all* of it, you'd have to stop the wind)
- Stronger winds → more power (like v^3)

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Chemical Energy



- Electrostatic energy (associated with charged particles, like electrons) is stored in the chemical bonds of substances.
- Rearranging these bonds can release energy (some reactions *require* energy to be put in)
- Typical numbers are 100–200 kJ per mole
 - a mole is 6.022×10^{23} molecules/particles
 - typical molecules are tens of grams per mole → works out to typical numbers like several thousand Joules per gram, or a few kilocalories per gram (remember, 1 kcal = 4184 J)

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Chemical Energy Examples

- Burning a wooden match releases about one Btu, or 1055 Joules (a match is about 0.3 grams), so this is $>3,000 \text{ J/g}$, nearly 1 kcal/g
- Burning coal releases about 20 kJ per gram of chemical energy, or roughly 5 kcal/g
- Burning gasoline yields about 39 kJ per gram, or just over 9 kcal/g
- Very few substances over about 11 kcal/g

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Energy from Food

- We get the energy to do the things we do out of food (stored solar energy in the form of chemical energy).
- Energy sources recognized by our digestive systems:
 - Carbohydrates: 4 kilocalories per gram
 - Proteins: 4 kilocalories per gram
 - Fats: 9 kilocalories per gram (like gasoline)

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Nutrition Labels

Whole Milk
Serving Size 8 fl oz (240mL)
Servings Per Container 2

Amount Per Serving
Calories 150 Calories from Fat 70

	% Daily Value*
Total Fat 8g	12%
Saturated Fat 5g	25%
Cholesterol 35mg	12%
Sodium 125mg	5%
Total Carbohydrate 12g	4%
Dietary Fiber 0g	0%
Sugars 11g	
Protein 8g	

Vitamin A 6% + Vitamin C 4%
Calcium 30% + Iron 0% + Vitamin D 25%
* Percent Daily Values are based on a 2,000 calorie diet. Your daily values may be higher or lower depending on your calorie needs.
Calories: 2,000 2,500

	Less than	50g	80g
Total Fat	Less than	25g	30mg
Sodium	Less than	2,400mg	300g
Total Carbohydrate	Less than	300g	375g
Dietary Fiber	Less than	25g	30g



- Nutrition labels tell you about the energy content of food
- Note that capital C means kcal
- Conversions: Fat: 9 kcal/g
Carbs: 4 kcal/g
Protein: 4 kcal/g
- This product has 72 kcal from fat, 48 kcal from carbohydrates, and 32 kcal from protein
 - sum is 152 kilocalories: compare to label
- 152 kcal = 636 kJ: enough to climb about 1000 meters (64 kg person)

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Our Human Energy Budget

- A 2000 kcal per day diet means $2000 \times 4184 \text{ J} = 8,368,000 \text{ J}$ per day
- 8.37 MJ in (24 hr/day) \times (60 min/hr) \times (60 sec/min) = 86,400 sec corresponds to 97 Watts of power
- Even a couch-potato at 1500 kcal/day burns 75 W
- More active lifestyles require greater caloric intake (more energy)





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Mass-energy

- Einstein's famous relation:
 $E = mc^2$
relates mass to energy
- In effect, they *are* the same thing
 - one can be transformed into the other
 - physicists speak generally of mass-energy
- Seldom experienced in daily life directly
 - Happens at large scale in the center of the sun, and in nuclear bombs and reactors
 - Actually *does* happen at barely detectable level in *all* energy transactions, but the effect is *tiny!*



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$E = mc^2$ Examples

- The energy equivalent of one gram of material (*any composition!!*) is $(0.001 \text{ kg}) \times (3.0 \times 10^8 \text{ m/s})^2 = 9.0 \times 10^{13} \text{ J} = 90,000,000,000,000 \text{ J} = 90 \text{ TJ}$
 - Man, that's big!
 - The U.S. energy budget is equivalent to 1000 kg/yr
- If one gram of material undergoes a *chemical* reaction, losing about 9,000 J of energy, how much *mass* does it lose?
 - $9,000 \text{ J} = \Delta mc^2$, so $\Delta m = 9,000/c^2 = 9 \times 10^3 / 9 \times 10^{16} = 10^{-13} \text{ kg}$ (would we *ever* notice?)

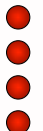
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Solar Energy is Nuclear, Using $E = mc^2$

- Thermonuclear fusion reactions in the sun's center
 - Sun is 16 million degrees Celsius in its center
 - Enough energy to ram protons together (despite mutual repulsion) and make deuterium, then helium
 - Reaction per atom 20 million times more energetic than chemical reactions, in general

4 protons:
mass = 4.029







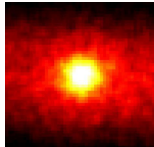
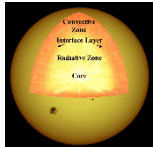
${}^4\text{He}$ nucleus:
mass = 4.0015

+ 2 neutrinos, photons (light)

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$E = mc^2$ in Sun

- Helium nucleus is *lighter* than the four protons!
- Mass difference is $4.029 - 4.0015 = 0.0276 \text{ a.m.u.}$
 - 1 a.m.u. (atomic mass unit) is $1.6605 \times 10^{-27} \text{ kg}$
 - difference of $4.58 \times 10^{-29} \text{ kg}$
 - multiply by c^2 to get $4.12 \times 10^{-12} \text{ J}$
 - 1 mole (6.022×10^{23} particles) of protons $\rightarrow 2.5 \times 10^{12} \text{ J}$
 - typical chemical reactions are 100-200 kJ/mole
 - **nuclear fusion is ~20 million times more potent stuff!**

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Energy from Light

- The tremendous energy from the sun is released as light. So light carries energy.
- How much??
- Best way to get at this is through the process of “blackbody” radiation, or thermal radiation...
- All objects emit “light”
 - Though almost all the light *we* see is *reflected* light
- The color and intensity of the emitted radiation depend on the object’s temperature

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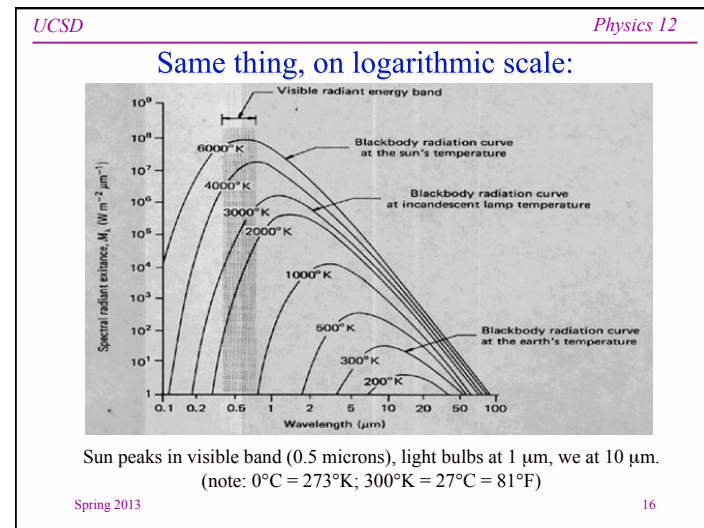
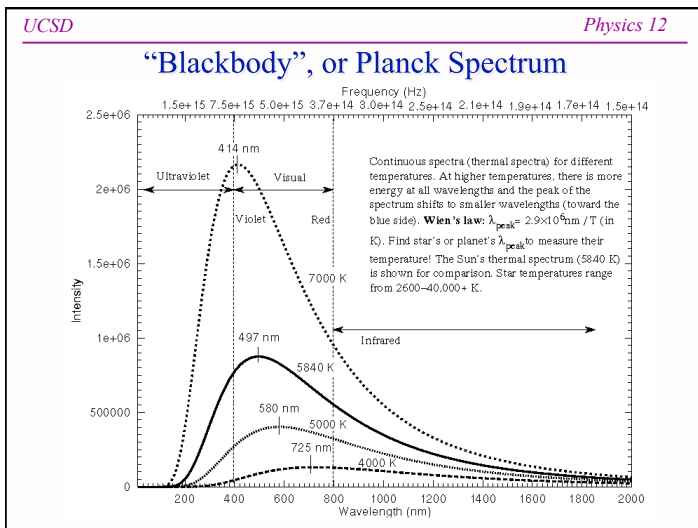
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Emitted Radiation’s Color and Intensity depend on Temperature

Object	Temperature	Color
You	~ 30 C	Infrared (invisible)
Heat Lamp	~ 500 C	Dull red
Candle Flame	~ 1700 C	Dim orange
Bulb Filament	~ 2700 C	Yellow
Sun’s Surface	~ 5500 C	Brilliant white

The hotter it gets, the “bluer” the emitted light
The hotter it gets, the *more intense* the radiation (more energy)

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Okay, but how much energy?

- The power given off of a surface in the form of light is proportional to the *fourth power* of temperature!
 - $F = \sigma T^4$ in Watts per square meter
 - the constant, σ , is numerically $5.67 \times 10^{-8} \text{ W}^\circ\text{K}^4/\text{m}^2$
 - easy to remember constant: 5678
 - temperature must be in Kelvin:
 - $^\circ\text{K} = ^\circ\text{C} + 273$
 - $^\circ\text{C} = (5/9) \times (^\circ\text{F} - 32)$
- Example: radiation from your body:
 - $(5.67 \times 10^{-8}) \times (310)^4 = 523$ Watts per square meter
 - (if naked in the cold of space: don't let this happen to you!)

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Radiant Energy, continued

- Example: The sun is 5800°K on its surface, so:
 - $F = \sigma T^4 = (5.67 \times 10^{-8}) \times (5800)^4 = 6.4 \times 10^7 \text{ W/m}^2$
 - Summing over entire surface area of sun gives $3.9 \times 10^{26} \text{ W}$
- Compare to total capacity of energy production on earth: $3.3 \times 10^{12} \text{ W}$
 - Single power plant typically 0.5–1.0 GW (10^9 W)
- In earthly situations, radiated power out partially balanced by radiated power in from other sources
 - Not 523 W/m^2 in 70°F room, more like 100 W/m^2
 - goes like $\sigma T_h^4 - \sigma T_c^4$

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And those are the major players...

- We've now seen all the major energy players we'll be discussing in this class:
 - work as force times distance
 - kinetic energy (wind, ocean currents)
 - gravitational potential energy (hydroelectric, tidal)
 - chemical energy (fossil fuels, batteries, food, biomass)
 - heat energy (power plants, space heating)
 - mass-energy (nuclear sources, sun's energy)
 - radiant energy (solar energy)

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Assignments

- Homework #1 due April 12 in class:
 - Chapter 1 problems, plus online additions
 - see assignments link on web page for details
 - don't forget to show your work/reasoning on the multiple choice!
 - the answers alone do not suffice
- Quiz #1 due Friday, April 12, by midnight
 - TED will be up and quizzes available by Thursday
 - 3 attempts permitted
 - all numerical/quantitative this week

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